Astronomy is rich with terminology. This column will help beginning stargazers ease into the world of astronomy by briefly introducing a new but basic astronomical term (word, acronym or abbreviation) each month. This list, which began January 1999 with the letter $\boldsymbol{a}$, is alphabetical but uses successive letters for each month's entry. (We will return to the letter $\boldsymbol{a}$ after twenty-six months.)

## Word of the Month for April 2000

parsec (symbol pc) A basic unit of stellar distance (like a mile or kilometer) used in professional astronomy to indicate stellar distances.

One parsec is the distance at which an object would have an annual heliocentric trigonometric parallax of one arc second. (An arc second, abbrev. arc sec or ", is $1 / 60$ of an arc minute or $1 / 3600$ of a degree.) See Figure 1 and end note.

This is equivalent to defining one parsec ( 1 pc ) as the distance at which the Earth's mean distance from the Sun (one astronomical unit or 1 AU ) would subtend an angle of one arc second (1'). See Figure 1.

One parsec is approximately:
3.2616 light years (ly)

206,265 astronomical units (AU)
30,857,000,000 kilometers (km)
19,174,000,000 miles (mi)
Multiples of the parsec include the kiloparsec (kpc, $1,000 \mathrm{pc}$ ) and the megaparsec ( $\mathrm{Mpc}, 1,000,000 \mathrm{pc}$ ).

Both the parsec and light year are commonly used to designate stellar distances. (The light year is the distance light travels through a vacuum in an average Gregorian year of 365.2425 days.)

The light year is often preferred in nontechnical work. The parsec is often used in professional astronomy because it is easily and conveniently related to the
parallax of a star:
Distance $($ in parsecs $)=1 /$ parallax (in arc seconds)
(Example: If the parallax $=0.5$ arc seconds, then the distance is 2 pc , which is about 6.5 ly .)

## Example Distances in Parsecs and Light Years

| Alpha Centauri | 1.3 pc | 4.2 ly |
| :--- | ---: | ---: |
| Sirius | 2.6 pc | 8.6 ly |
| Vega | 7.7 pc | 25 ly |
| Regulus | 24 pc | 77 ly |
| Antares | 60 pc | 197 ly |
| Pleiades Star Cluster | 126 pc | 410 ly |
| Polaris | 132 pc | 431 ly |
| Betelgeuse | 160 pc | 522 ly |
| Deneb | 450 pc | $1,467 \mathrm{ly}$ |
| Orion Nebula | 460 pc | $1,500 \mathrm{ly}$ |
| Hercules Globular Cluster | 640 pc | $21,000 \mathrm{ly}$ |

Fig. 1. Parallax. If parallax $=1$ arc $s e c$, distance $=1 \mathrm{pc}$


## Note:

Parallax is the apparent change in the direction of an object, caused by a change in the observer's position.

The annual heliocentric parallax results from viewing the angular shift of an object by moving through a distance equal to the mean distance between the Earth and Sun (called one astronomical unit or 1 AU ).

References. J. Mitton 1991, Concise Dictionary of Astronomy (Oxford Univ. Press); I. Ridpath 1997, A

Dictionary of Astronomy (Oxford Univ. Press).

